

# X-ray spectroscopy - a walk through XSPEC

- Download data:
  - go to [http://galaxy.asu.cas.cz/~svoboda/pub/AGN\\_workshop/data/](http://galaxy.asu.cas.cz/~svoboda/pub/AGN_workshop/data/) and download folders spectrum\_1 and spectrum\_2
- Load data in the Xspec
  - go to the folder Spectrum\_1 and load EPIC/PN data:
    - **data** mcg6\_10\_r.pi
  - check what data you have:
    - **show data**
  - ignore bad data and data out of valid energy range:
    - **ignore** bad
    - **ignore** \*\*-1. 10.-\*\* (opposite command is **notice**)
      - \*\*- means everything up to., -\*\* means everything from
      - be aware that if you put integers behind ignore/notice, it will not ignore energies but channels!
    - **notice** 0.5-1. (XMM-Newton is good also below 1 keV)
      - if you have more data and you want to apply notice/ignore to all of them, just type "\*\*:": e.g., ignore \*\*:\*\*-0.5 10.-\*\*
- Plot your data
  - **plot ldata** (plot data, plot counts)
- Define a model
  - **model phabs\*powerlaw**
    - enter model parameters
  - change a parameter of the model, freeze known parameters
    - **newpar** 1 0.036
    - **freeze** 1 (opposite command is **thaw**)
  - you can also define delta of a parameter, which is a first step of the value for the fitting procedure (if negative the parameter is frozen), you can also define allowed intervals for the parameter: newpar #param\_number #param\_value, #delta, #min, #min, #max, #max
    - e.g. newpar 2 2.,0.1,1.5,1.5,3.,3.
  - show free/all parameters
    - e.g., **show free**
    - **show all**
- Fit the data
  - **fit**
  - if you want to define number of steps in fitting and the depth (critical delta), you can type e.g.:
    - fit 1000 1e-2
  - plot the data and model residuals, they are different ways of plotting residuals:
    - **plot ldata ratio**
    - **plot ldata residuals**
    - plot ldata chisq, ...
- Add a component to your model
  - **addc** 2 zxipcf
  - addc 3 zgau
  - fit with the new model, has the fit improved?
- Save your model/all
  - **save model** my\_model.mdl
  - **save all** my\_session.xcm
- Load model/all:

- @my\_model.mdl
  - @my\_session.xcm
- Check your model if it makes sense:
  - what is the energy and sigma for the gaussian line, any physical interpretation?
    - plot model
- Replace a Gaussian line by a relativistic iron line:
  - delete a model component:
    - **delc 3**
  - add a new component, you can also use "editmod" instead of "addc"
    - e.g., **editmod** phabs\*zxipcf\*(po+kyrline)
- Try to fit, change parameters if needed. Try to get the least residuals as possible and achieve the reduced chi-2 less than two (reduced chi-2 is chi-2 divided by degrees of freedom).
- Calculate the error of a parameter (must be a free parameter):
  - **error 6**
  - error 8
  - error command works only if your reduced chi-2 is less than 2; if it is higher and you want to run this command, you can do it by typing, e.g., error max 10 8
- Calculate the dependence of chi-2 on the parameter value:
  - **steppar 8 0 1 100**
  - **plot contour**
- Calculate a 2D contour between two parameters:
  - if you want Xspec to start calculation in each grid from the best fit type "steppar best ..."
    - steppar best 8 0 1 20 9 20 45 20
    - plot contour
- Write your results in a file
  - **setplot command** we myfit
  - plot ldata ratio (or what you want to write in your file)
  - **setplot delete 1** (otherwise your file will be rewritten again by new commands)
- Save all and exit:
  - save all bestfit.xcm
  - exit
- **Good luck with fitting!**

## Additional Exercise:

- 1) Investigate two unknown X-ray spectra and find out which is Type-1 and which is Type-2 AGN:

Spectrum 1:  $z = 0.008$ , Galactic column density  $nH = 3.6e20 \text{ cm}^{-2}$

Spectrum 2:  $z = 0.015$ , Galactic column density  $nH = 5.7e20 \text{ cm}^{-2}$

Cross the line between the corresponding AGN spectrum and type:

AGN spectrum 1

Seyfert-1 AGN

AGN spectrum 2

Seyfert-2 AGN

- 2) Fit the spectra with a simple absorbed power-law model and measure the power-law slope and the level of absorption. Add a reflection model (line "gauss" and/or reflection continuum "pexrav") to your spectral fit and check for the presence of reflection features. If present measure the reflection fraction or the equivalent width of the iron line.
- 3) Calculate uncertainty of the interesting parameters and plot a statistical contour (in 1 or 2-D parameter space).
- 4) Advanced exercise: in Spectrum 1 load a spectrum from a different time interval (in folder "Additional") and compare the two spectra from different time intervals. Which spectrum was caught in high/low state and what changed in the spectrum?